



• For this to give the proper average, *f* is normalized as follows:

 $\int d\mathbf{r} \int d\mathbf{v} f(\mathbf{v}, \mathbf{r}, t) = 1$ 

 To derive an equation of motion for f(v,r,t), it is somewhat easier to consider the particle density

$$n(\mathbf{v},\mathbf{r},t) = Nf(\mathbf{v},r,t)$$

where

$$N = \int d\mathbf{r} \int d\mathbf{v} n(\mathbf{v}, \mathbf{r}, t) = Total \ \# of \ particles$$

• The density *n*(**v**,**r**,*t*) should satisfy a continuity equation in the 6D phase space defined by

 $x, y, z, v_x, v_y, v_z \rightarrow$  Independent variables



• The flux density is written in terms of the time derivatives of  
the 'position' variables in 6D:  
$$j(x, y, z, v_x, v_y, v_z) = v_x n(v, r, t) \hat{a}_x + v_y n \hat{a}_y + v_z n \hat{a}_z + \frac{F_x}{m} n \hat{b}_{v_x} + \frac{F_y}{m} n \hat{b}_{v_y} + \frac{F_z}{m} n \hat{b}_{v_z}$$
 with  $\dot{v} = \frac{F}{m}$   
• Applying the divergence theorem in 6D  
$$\int_S j(v, r, t) \cdot ds = \int_V dr dv \nabla \cdot j(v, r, t)$$
where the divergence of **j** is  
$$\nabla \cdot j = v_x \frac{\partial n}{\partial x} + v_y \frac{\partial n}{\partial y} + v_z \frac{\partial n}{\partial z} + \frac{F_x}{m} \frac{\partial n}{\partial v_x} + \frac{F_y}{m} \frac{\partial n}{\partial v_y} + \frac{F_z}{m} \frac{\partial n}{\partial v_z}$$

which is written more compactly as:  

$$\nabla \cdot \mathbf{j} = \mathbf{v} \cdot \nabla_r n + \frac{\mathbf{F}}{m} \nabla_v n$$
• Particle balance is therefore:  

$$\int_{V} d\mathbf{r} d\mathbf{v} \left( \frac{\partial n}{\partial t} + \mathbf{v} \cdot \nabla_r n + \frac{\mathbf{F}}{m} \cdot \nabla_v n - \frac{\partial n}{\partial t} \Big|_{Coll} - \frac{\partial n}{\partial t} \Big|_{G-R} \right) = 0$$
Normalizing, we get the classical form of the Boltzmann transport equation:  

$$\frac{\partial f(\mathbf{r}, \mathbf{v}, t)}{\partial t} = -\mathbf{v} \cdot \nabla_r f - \frac{\mathbf{F}}{m} \cdot \nabla_v f + \frac{\partial f}{\partial t} \Big|_{Coll} + \frac{\partial f}{\partial t} \Big|_{G-R}$$
First two terms on the rhs are the streaming terms

 For Bloch electrons in a semiconductor, we could have considered a 6D space x,y,z,k<sub>x</sub>,k<sub>y</sub>,k<sub>z</sub> where k is the wavevector and

$$\mathbf{v} = \frac{1}{\hbar} \nabla_k \boldsymbol{E}(\mathbf{k})$$

• The semi-classical BTE for transport of Bloch electrons is therefore

$$\frac{\partial f(\mathbf{r},\mathbf{k},t)}{\partial t} = -\frac{1}{\hbar} \nabla_{k} \boldsymbol{E}(\mathbf{k}) \cdot \nabla_{r} f - \frac{\mathbf{F}}{\hbar} \cdot \nabla_{k} f + \frac{\partial f}{\partial t} \Big|_{Coll} + \frac{\partial f}{\partial t} \Big|_{G-R}$$



(A) Out Scattering

$$\Delta n(\mathbf{r},\mathbf{k},t) = -n(\mathbf{r},\mathbf{k},t)\Gamma_{\mathbf{k}\mathbf{k}'}\Delta t$$

where  $\Gamma_{\! kk'}$  is the transition rate per particle from  ${\bm k}$  to  ${\bm k'}$ 

Distribution function is:  $f(\mathbf{r}, \mathbf{k}, t) = \frac{n(\mathbf{r}, \mathbf{k}, t)}{N}$ 

Take limit as  $\Delta t \rightarrow 0$ 

$$\frac{\partial f(\mathbf{r},\mathbf{k},t)}{\partial t}\Big|_{OUT} = -f(\mathbf{r},\mathbf{k},t)\Gamma_{\mathbf{k}\mathbf{k}'}[1-f(\mathbf{r},\mathbf{k}',t)]$$

where the last term in brackets accounts for the Pauli exclusions principle (degeneracy of the final state after scattering).

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(B) In ScatteringBy an analogous argument, the rate of change of the distribution function due to in scattering is:  $\frac{\partial f(\mathbf{r}, \mathbf{k}, t)}{\partial t} \Big|_{IN} = f(\mathbf{r}, \mathbf{k}', t)\Gamma_{\mathbf{k}'\mathbf{k}}[1 - f(\mathbf{r}, \mathbf{k}, t)]$ Total rate of change of  $f(\mathbf{r}, \mathbf{k}, t)$  around  $\mathbf{k}$  is a sum over all possible initial and final states  $\mathbf{k}'$ :  $\frac{\partial f(\mathbf{r}, \mathbf{k}, t)}{\partial t} \Big|_{Coll} = \sum_{\mathbf{k}'} \{f(\mathbf{r}, \mathbf{k}', t)[1 - f(\mathbf{r}, \mathbf{k}, t)]\Gamma_{\mathbf{k}'\mathbf{k}} - f(\mathbf{r}, \mathbf{k}, t)[1 - f(\mathbf{r}, \mathbf{k}, t)]\Gamma_{\mathbf{k}\mathbf{k}'}\}$ Out scattering (C) Boltzmann Equation with Collision Integral

The sum over final states  $\mathbf{k}'$  may be converted to an integral due to the small volume of k-space associated with each state:

$$\sum_{\mathbf{k}'} \rightarrow \frac{V}{8\pi^3} \int d\mathbf{k}'$$

The BTE becomes:

$$\frac{\partial f_{k}}{\partial t} + \frac{1}{\hbar} \nabla_{k} \boldsymbol{E} \cdot \nabla_{r} f_{k} + \frac{F}{\hbar} \nabla_{k} f_{k} = \frac{V}{8\pi^{3}} \int d\mathbf{k} \{ f_{k'} [1 - f_{k}] \Gamma_{\mathbf{k'k}} - f_{k} [1 - f_{k'}] \Gamma_{\mathbf{kk'}} \}$$



#### Time dependent perturbation theory

• Assume the Hamiltonian may be decomposed as  $H=H_0+V_s$ , where  $H_0$  is the Hamiltonian of the perfect crystal (described by Bloch states),  $V_s(\mathbf{r},t)$  is a small random potential. If  $V_s << H_0$ , then it is a good approximation to expand the solution (with random part) in terms of unperturbed eigenstates:

$$H_0 \Psi_k = E_k \Psi_k; \quad \Psi_k^0(\mathbf{r}, t) = \Psi_k(\mathbf{r}) e^{-iE_k t/\hbar}$$

• Expand actual solution in terms of these orthonormal functions:

$$\Psi(\mathbf{r},t) = \sum_{k} c_{k}(t) \Psi_{k}(\mathbf{r}) e^{-iE_{k}t/\hbar}$$





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• Assume sufficiently weak scattering that  $c_{ko} \approx 1$ , and  $c_{k \neq ko} \approx 0$  for all time. The dominant term in the sum is:

$$i\hbar \frac{\partial \boldsymbol{c}_{\boldsymbol{k}_{0}}(t)}{\partial t} = \boldsymbol{c}_{\boldsymbol{k}_{0}}(t) \langle \boldsymbol{k}_{0}' | \boldsymbol{V}_{\boldsymbol{s}} | \boldsymbol{k}_{0} \rangle \boldsymbol{e}^{-i(\boldsymbol{E}_{\boldsymbol{k}_{0}} - \boldsymbol{E}_{\boldsymbol{k}_{0}})t/\hbar}$$

which integrates to

$$c_{k_0}(t) = \frac{1}{\hbar} \int_0^t dt' \langle k_0' | V_s | k_0 \rangle e^{-i(E_{k_0} - E_{k_0})t'/\hbar} + c_{k_0'}(0)$$

• Suppose  $V(\mathbf{r},t)$  may be Fourier decomposed, so that

$$V_s(\mathbf{r},t) = V_s(\mathbf{r}) e^{\pm i\omega t}$$

Note that this form of  $V(\mathbf{r},t)$  may correspond to interaction with lattice vibrations or with optical excitation.

Then substituting

$$\boldsymbol{C}_{\boldsymbol{K}_{0}^{\prime}}(t) = \frac{1}{i\hbar} \langle \boldsymbol{K}_{0}^{\prime} | \boldsymbol{V}_{s} | \boldsymbol{K}_{0} \rangle_{0}^{t} \boldsymbol{d}t^{\prime} \boldsymbol{e}^{-i\Lambda t^{\prime}}; \ \Lambda = \left(\boldsymbol{E}_{\boldsymbol{K}_{0}^{\prime}} - \boldsymbol{E}_{\boldsymbol{K}_{0}} \mp \hbar \boldsymbol{\omega}\right) / \hbar$$

and integrating this last expression leads to

$$c_{k_0'}(t) = \frac{1}{i\hbar} V^{k_0 k_0'} s \frac{e^{-i\Lambda t} - 1}{i\Lambda}$$
$$c_{k_0'}(t) = \frac{1}{i\hbar} V^{k_0 k_0'} s e^{-i\Lambda t/2} \left(\frac{\sin(\Lambda t)}{\Lambda t}\right) t$$

• Since the probability of being in  $\mathbf{k}_0'$  is given by

$$P_{k_0k_0'} = \lim_{t \to \infty} \left| c_{k_0'}(t) \right|^2$$

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• Substituting for *c* and taking the magnitude squared gives  $P_{k_0k'_0} = \lim_{t \to \infty} \frac{1}{\hbar^2} |V_s^{k_0k'_0}|^2 \left(\frac{\sin(\Lambda t)}{\Lambda t}\right)^2 t^2$ where asymptotically  $\lim_{t \to \infty} \left(\frac{\sin(\Lambda t)}{\Lambda t}\right)^2 = 2\pi\delta(\Lambda)/t = 2\pi\hbar\delta(E_{k'_0} - E_{k_0} \mp \hbar\omega)/t$ This gives the famous **Fermi's Golden Rule** (droping 0s index)  $\Gamma_{kk'} = \frac{P_{kk'}}{t} = \frac{2\pi}{\hbar} |V_s^{kk'}|^2 \delta(E_{k'} - E_k \mp \hbar\omega)$ • Assumptions made: (1) Long time between scattering (no multiple scattering events) (2) Neglect contribution of other *c*'s (Collision broadening ignored) **Computational Electronics** 





• Considering the induced charge caused by the change in the electron gas by the impurity, the net potential seen is

$$V_i(\mathbf{q}) = \frac{V_i^{0}(\mathbf{q})}{\varepsilon(\mathbf{q},\omega)}$$

In the above expression, **q** is the wavevector associated with Fourier transforming the potential (and Poisson's equation),  $V_i(\mathbf{q})$  is the total potential seen by an electron due to an impurity, and  $\varepsilon(\mathbf{q}, \omega)$  is the *dielectric function* characterizing the polarization of the electron gas to the impurity potential.

 In linear response theory, this may be calculated in the random phase approximation (RPA) to give the Lindhard dielectric function

$$\varepsilon(\mathbf{q},\boldsymbol{\omega}) = 1 - \lim_{s \to \infty} \frac{\boldsymbol{e}^2}{\varepsilon_{sc} \boldsymbol{q}^2} \sum_{\mathbf{k}} \frac{f_0(\boldsymbol{E}_{\mathbf{k}+\mathbf{q}}) - f_0(\boldsymbol{E}_{\mathbf{k}})}{\boldsymbol{E}_{\mathbf{k}+\mathbf{q}} - \boldsymbol{E}_{\mathbf{k}} + \hbar\boldsymbol{\omega} + i\delta}$$

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• Assuming low frequencies, and assuming long wavelengths, the Thomas-Fermi function is obtained to be of the form:

$$\lim_{\omega,q\to 0} \varepsilon(q,\omega) \approx 1 + \frac{\lambda^2}{q^2}$$

where the inverse screening length  $\lambda^2$  is given as (3D):

$$\lambda^2 = \frac{e^2 n}{\varepsilon_{sc} k_B T}$$
 high temperature;  $\lambda^2 = \frac{3e^2 n}{2\varepsilon_{sc} E_F}$ ;  $T = 0K$ 

In here, *n* is the carrier density and  $E_F$  is the Fermi energy.

• Assuming the Fermi Thomas form, inverse Fourier transforming gives the form of the screened potential in real space as:

$$V_i(\mathbf{r}) = -\frac{Zq^2}{4\pi\varepsilon r}e^{-\lambda t}$$

• For the scattering rate due to impurities, we need for Fermi's rule the matrix element between initial and final Bloch states

 $\langle n', \mathbf{k}' | V_i(\mathbf{r}) | n, \mathbf{k} \rangle = V^{-1} \int d\mathbf{r} u_{n',k'}^* e^{-i\mathbf{k}' \cdot \mathbf{r}} V_i(\mathbf{r}) u_{n,k} e^{i\mathbf{k} \cdot \mathbf{r}}$ 

Since the *u*'s have periodicity of lattice, expand in reciprical space

$$=\sum_{G} V^{-1} \int d\mathbf{r} e^{-i\mathbf{k}'\cdot\mathbf{r}} V_{i}(\mathbf{r}) e^{i\mathbf{k}\cdot\mathbf{r}} e^{-i\mathbf{G}\cdot\mathbf{r}} U_{nn'kk'}(\mathbf{G})$$
  
$$=\sum_{G}^{G} V^{-1} \int d\mathbf{r} e^{-i\mathbf{k}'\cdot\mathbf{r}} V_{i}(\mathbf{r}) e^{i\mathbf{k}\cdot\mathbf{r}} e^{-i\mathbf{G}\cdot\mathbf{r}} \int_{\Omega} d\mathbf{r}' u_{n',k'}^{*}(\mathbf{r}') u_{n,k}(\mathbf{r}') e^{i\mathbf{G}\cdot\mathbf{r}'}$$

• For impurity scattering, the matrix element has a 1/q type dependence which usually means **G**≠0 terms are small

$$= V^{-1} \int d\mathbf{r} e^{-i\mathbf{k}' \cdot \mathbf{r}} V_i(\mathbf{r}) e^{i\mathbf{k} \cdot \mathbf{r}} \int_{\Omega} d\mathbf{r}' u_{n',k'}^*(\mathbf{r}') u_{n,k}(\mathbf{r}') = V_i(\mathbf{q}) I_{kk'}$$

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 The usual argument is that since the u's are normalized within a unit cell (i.e. equal to 1), the Bloch overlap integral *I*, is approximately 1 for n'=n [interband(valley)]. Therefore, for impurity scattering, the matrix element for scattering is approximately

$$|\langle \mathbf{k}' | V_i(\mathbf{r}) | \mathbf{k} \rangle|^2 = |V_i(\mathbf{q})|^2 \cong \frac{Z^2 e^4}{V^2 (q^2 + \lambda^2) \epsilon_{sc}^2}; V = volume$$

where the scattered wavevector is:  $\boldsymbol{q} = \boldsymbol{k} - \boldsymbol{k}'$ 

• This is the scattering rate for a single impurity. If we assume that there are *N<sub>i</sub>* impurities in the whole crystal, and that scattering is completely uncorrelated between impurities:

$$V_i^{kk'} \cong \frac{N_i Z^2 e^4}{V^2 (q^2 + \lambda^2) \epsilon_{sc}^2} = \frac{n_i Z^2 e^4}{V (q^2 + \lambda^2) \epsilon_{sc}^2}$$

where  $n_i$  is the impurity density (per unit volume).

 The total scattering rate from k to k' is given from Fermi's golden rule as:

$$\Gamma_{kk'}^{i} = \frac{2\pi n_{i}Z^{2}e^{4}}{V\hbar(q^{2}+\lambda^{2})\varepsilon_{sc}^{2}}\delta(E_{k'}-E_{k})$$

If  $\theta$  is the angle between **k** and **k**', then:

$$\boldsymbol{q} = |\mathbf{k} - \mathbf{k}'| = \boldsymbol{k}^2 + \boldsymbol{k}'^2 - 2\boldsymbol{k}\boldsymbol{k}'\cos\theta = 2\boldsymbol{k}^2(1 - \cos\theta)$$

• Comments on the behavior of this scattering mechanism:

- Increases linearly with impurity concentration
- Decreases with increasing energy  $(k^2)$ , favors lower T
- Favors small angle scattering
- Ionized Impurity-Dominates at low temperature, or room temperature in impure samples (highly doped regions)
- Integration over all **k**' gives the total scattering rate  $\Gamma_k$ :

$$\Gamma_{k}^{i} = \frac{n_{i}Z^{2}e^{4}m^{*}}{8\pi\varepsilon_{sc}^{2}\hbar^{3}k^{3}} \left[\frac{4k^{2}}{q_{D}^{2}(4k^{2}+q_{D}^{2})}\right]; \quad q_{D} = 1/\lambda$$













### • Phonons:

The Fourier expansion in reciprocal space of the coupled vibrational motion of the lattice decouples into *normal modes* which look like an independent set of Harmonic oscillators with frequency  $\omega_{q}^{\xi}$ 

 $\xi$  labels the mode index, acoustic (longitudinal, 2 transverse modes) or optical (1 longitudinal, 2 transverse)

 ${\bf q}$  labels the wavevector corresponding to traveling wave solutions for individual components,

• The velocity and the occupancy of a given mode are given by:

$$v_{q}^{\xi} = \frac{\partial \omega_{q}^{\xi}}{\partial q}$$

$$n_{q}^{\xi} = \frac{1}{e^{\hbar \omega_{q}^{\xi}/k_{B}T_{i}} - 1}; \quad Bose - Einstein \ distribution$$
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where:  $\Delta = \nabla \cdot \mathbf{u}(\mathbf{r}) = \text{dilation of volume of unit cell}$   $E_1 = \text{Deformation potential const.}$   $E_1\Delta = \text{Deformation potential}$ and  $\mathbf{u}$  is the displacement operator of the lattice  $\mathbf{u}^{\xi}(\mathbf{r}) = \sum_{q} \left(\frac{\hbar}{2NM\omega_q^{\xi}}\right)^{1/2} \vec{e}_{q,\xi} \left(a_{q,\xi} e^{i\mathbf{q}\cdot\mathbf{r}} + a_{q,\xi}^{*} e^{-i\mathbf{q}\cdot\mathbf{r}}\right)$   $\vec{e}_{q,\xi} = \text{polarization vector}$ • Taking the divergence gives factor of  $\mathbf{e}\cdot\mathbf{q}$  of the form:  $\vec{e}_{q,\xi} \cdot \mathbf{q} = q \text{ for longitudin al modes}$   $\vec{e}_{q,\xi} \cdot \mathbf{q} = 0 \text{ for transverse modes}$ Therefore, only longitudinal modes contribute. • For ellipsoidal valleys (i.e. Si, Ge), shear strains may contribute to the scattering potential

$$E(\mathbf{k}) \cong E^{0}(\mathbf{k}) + E_{d}\Delta + E_{u}e_{zz}$$
$$e_{zz} \cong \frac{\partial u}{\partial z} \cdot \hat{z}; \quad e_{zz} \text{ is component of the strain tensor}$$

**Scattering Matrix Element:** 

Assuming  $\omega_q = u_I q$ , then:

$$|V_{ac}|^{2} = \frac{\hbar E_{1}^{2} q_{\pm} (n_{q} + 1 \mp 1)}{2V \rho u_{l}} \begin{pmatrix} upper \ absorption \\ lower \ emission \end{pmatrix}$$

• At sufficient high temperature, (equipartition approximation):

$$n_{\rm q} \approx n_{\rm q} + 1 \approx \frac{k_B T_I}{\hbar \omega_{\rm q}}$$

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• Substituting and assuming linear dispersion relation, Fermi's rule becomes

$$\Gamma_{kk'}^{ac} = \frac{2\pi}{\hbar} |V_{ac}|^2 \delta \left( E_{k'} - E_k \mp \hbar \omega_q \right) = \frac{2\pi E_1^2 k_B T_l}{\hbar V \rho u_l^2} \delta \left( E_{k'} - E_k \mp \hbar \omega_q \right)$$

• The total scattering rate due to acoustic modes is found by integrating over all possible final states **k**'

$$\Gamma_{k}^{ac} = \frac{2\pi E_{1}^{2} k_{B} T_{I}}{\hbar V \rho u_{I}^{2}} \frac{V}{8\pi^{3}} (4\pi) \int_{0}^{\infty} d\mathbf{k}' \mathbf{k}'^{2} \delta \left( E_{\mathbf{k}'} - E_{\mathbf{k}} \mp \hbar \omega_{\mathbf{q}} \right)$$

where the integral over the polar and azimuthal angles just gives  $4\pi$ .

• For acoustic modes, the phonon energies are relatively small since

 $\omega_{q} \rightarrow 0$  as  $q \rightarrow 0$ 

• Integrating gives (assuming a parabolic band model)

$$\Gamma_{k}^{ac} = \frac{m^{*} k E_{1}^{2} (k_{B} T)_{l}}{\pi \hbar^{3} c_{l}}; \ c_{l} = \rho u_{l}^{2}$$

where  $c_i$  is the longitudinal elastic constant. Replacing *k*, using parabolic band approximation, finally leads to:

$$\Gamma_{k}^{ac} = \frac{\sqrt{2}m^{*3/2}E_{1}^{2}(k_{B}T)_{l}}{\pi\hbar^{4}c_{l}}E^{1/2}$$

• Assumptions made in these derivations:

- a) spherical parabolic bands
- b) equipartition (not valid at low temperatires)
- c) quasi-elastic process (non-dissipative)

d) deformation potential Ansatz

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(B) Optical deformation potential scattering

(Due to symmetry of CB states, forbidden for  $\Gamma$ -minimas)

• Assume no dispersion:

$$\omega_{_{
m d}} 
ightarrow \omega_{_0}$$
 as  $q 
ightarrow 0$ 

Out of phase motion of basis atoms creates a strain called the optical strain.

• This takes the form ( $D_0$  is optical deformation potential field)

$$V_{do} = \vec{D}_0 \cdot u(\mathbf{r}); \quad D_0 \approx \vec{D}_0 \cdot \vec{e}_q \quad zeroth \ order$$

The matrix element for spherical bands is given by

$$\left|V_{kk'}^{ac}\right|^{2} = \left(\frac{\hbar D_{0}^{2}}{2\rho V \omega_{0}}\right) \left[n_{\omega_{0}} \delta(\mathbf{k} - \mathbf{k'} + \mathbf{q}) + \left[n_{\omega_{0}} + 1\right] \delta(\mathbf{k} - \mathbf{k'} - \mathbf{q})\right]$$

which is independent of q.

• The total scattering rate is obtained by integrating over all **k**' for both absorption and emission

$$\Gamma_{k}^{do} = \frac{1}{\tau_{do}} = \frac{m^{*3/2} D_{0}^{2}}{\sqrt{2} \pi \rho \hbar^{3} \omega_{0}} \begin{cases} n_{\omega_{0}} (E + \hbar \omega_{0})^{1/2} + [n_{\omega_{0}} + 1] (E - \hbar \omega_{0})^{1/2} \Theta (E - \hbar \omega_{0}) \end{cases}$$

where the first term in brackets is the contribution due to absorption and the second term is that due to emission

- For non-spherical valleys, replace  $m_t^{*^{3/2}} \Rightarrow m_t m_l^{1/2}$
- The non-polar scattering rate is basically proportional to density of states

$$\Gamma_k^{do} \propto \rho(E \pm \hbar \omega_0)$$



 Since Δk is large compared to k, assume ω<sub>q</sub> → ω<sub>Δq</sub> and treat the scattering the same as non-polar optical scattering replacing D<sub>0</sub> with D<sub>ij</sub> the intervalley deformation potential field, and the phonon coupling valleys *i* and *j*

$$\omega_{q} \rightarrow \omega_{ii}$$

• Conservation of energy also requires that the difference in initial and final valley energy be accounted for, giving

$$\Gamma_{k}^{i\nu} = \sum_{j} \frac{m_{d_{j}}^{3/2} D_{jj}^{2}}{\sqrt{2}\pi \rho \hbar^{3} \omega_{ij}} \begin{cases} n_{\omega_{ij}} (E - \Delta E_{ij} + \hbar \omega_{ij})^{1/2} + [n_{\omega_{ij}} + 1] (E - \Delta E_{ij} - \hbar \omega_{ij})^{1/2} \Theta (E - \Delta E_{ij} - \hbar \omega_{0}) \end{cases}$$

where the sum is over all the final valleys, j and

$$\Delta \boldsymbol{E}_{ij} = \boldsymbol{E}_{\min j} - \boldsymbol{E}_{\min j}$$









• The longitudinal displacement of the two atoms in the unit cell leads to a polarization dipole:

$$\mathbf{P} = \frac{N}{2V} e^* \mathbf{w} = \frac{N e^{*2} / 2V \overline{M}}{\omega_{TO}^2 - \omega^2} \mathbf{E}$$

• The existence of a finite polarization dipole modifies the dielectric function:

$$\mathbf{D} = \varepsilon_{\infty} \mathbf{E} + \mathbf{P} = \varepsilon_{\infty} \mathbf{E} + \frac{N e^{*2} / 2V \overline{M}}{\omega_{TO}^2 - \omega^2} \mathbf{E} = \varepsilon(\omega) \mathbf{E}$$
$$\varepsilon(\omega) = \varepsilon_{\infty} \left( 1 + \frac{S}{\omega_{TO}^2 - \omega^2} \right) = \varepsilon_{\infty} \left( 1 + \frac{\omega_{LO}^2 - \omega_{TO}^2}{\omega_{TO}^2 - \omega^2} \right)$$
$$Polarization \longrightarrow S = \frac{N e^{*2}}{2V \varepsilon_{\infty} \overline{M}} = \omega_{LO}^2 \varepsilon_{\infty} \left( \frac{1}{\varepsilon_{\infty}} - \frac{1}{\varepsilon(0)} \right)$$
$$S = Computational Electronics$$

• The electric field associated with the perturbed dipole moment is obtained from the condition that, in the absence of macroscopic free charge:  

$$\nabla \cdot \mathbf{D} = i\mathbf{k} \cdot \mathbf{D} = \nabla \cdot (\varepsilon_{\infty} \mathbf{E}_{ind} + \mathbf{P}) = 0 \Rightarrow \mathbf{D} = 0 \text{ and } \mathbf{E}_{ind} = -\frac{\mathbf{P}}{\varepsilon_{\infty}}$$
• Consider only one Fourier component:  

$$\mathbf{E}_{ind} = -\nabla \phi_{ind} = \frac{1}{e} \nabla V(q) = \frac{1}{e} \nabla \left[ V_{\mathbf{q}} e^{i\mathbf{q}\cdot\mathbf{r}} \right] = i \frac{1}{e} \mathbf{q} V(q)$$
•  $V(q) = i \frac{e}{\varepsilon_{\infty} q^2} \mathbf{q} \cdot \mathbf{P} \Rightarrow V(q) = i \frac{e}{\varepsilon_{\infty} q} P$ 
•  $V(\mathbf{r}) = i \frac{e}{\varepsilon_{\infty}} \frac{N}{2V} e^* \sqrt{\frac{\hbar}{2\overline{M}(N/2)\omega_{LO}}} \sum_{q} \frac{1}{q} (a_{\mathbf{q}} e^{i\mathbf{q}\cdot\mathbf{r}} + a_{\mathbf{q}}^+ e^{-i\mathbf{q}\cdot\mathbf{r}})$ 

$$= i \left(\frac{\hbar e^2}{2V\gamma\omega_{LO}}\right)^{1/2} \sum_{q} \frac{1}{q} (a_{\mathbf{q}} e^{i\mathbf{q}\cdot\mathbf{r}} + a_{\mathbf{q}}^+ e^{-i\mathbf{q}\cdot\mathbf{r}}), \quad \frac{1}{\gamma} = \omega_{LO}^2 \left[\frac{1}{\varepsilon_{\infty}} - \frac{1}{\varepsilon(0)}\right]$$
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# **Scattering Rate Calculation:**

• Matrix element squared for this interaction:

$$\left|V_{k,k'}\right|^{2} = \frac{\hbar \boldsymbol{e}^{2}}{2V\gamma\omega_{LO}}\frac{1}{\boldsymbol{q}^{2}}\left(\boldsymbol{N}_{0} + \frac{1}{2}\mp\frac{1}{2}\right)\delta(\boldsymbol{k'}-\boldsymbol{k}\mp\boldsymbol{q})$$

• Transition rate per unit time from state **k** to state **k**':

$$\Gamma_{k,k'} = \frac{2\pi}{\hbar} |V_{k,k'}|^2 \delta(\boldsymbol{\varepsilon}_{\mathbf{k}'} - \boldsymbol{\varepsilon}_{\mathbf{k}} \mp \hbar \boldsymbol{\omega}_{L0})$$
  
=  $\frac{\pi e^2}{V \gamma \omega_{LO}} \frac{1}{q^2} (N_0 + \frac{1}{2} \mp \frac{1}{2}) \delta(\mathbf{k'} - \mathbf{k} \mp \mathbf{q}) \delta(\boldsymbol{\varepsilon}_{\mathbf{k}'} - \boldsymbol{\varepsilon}_{\mathbf{k}} \mp \hbar \boldsymbol{\omega}_{L0})$ 

• Total scattering rate per unit time out of state k:

$$\Gamma_{k} = \sum_{k'} \Gamma_{k,k'} = \sum_{q} \Gamma_{k,q} = \frac{V}{(2\pi)^{3}} \int_{0}^{2\pi} d\varphi \int_{-1}^{1} d(\cos \theta) \int_{0}^{\infty} q^{2} \Gamma_{k,q} dq$$
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## (B) Piezoelectric scattering

• Since the polarization is proportional to the acoustic strain, we have

$$\mathbf{P} = \mathbf{e}_{pz} \nabla \cdot \mathbf{u}$$

• Following the same arguments as for the polar optical phonon scattering, one finds that the matrix element squared for this mechanism is:

$$|V_{kk'}|^2 = \frac{\hbar}{2\rho V \omega_{qv}} \left(\frac{e e_{\rho z}}{\varepsilon_{\infty}}\right)^2 \left(N_q + \frac{1}{2} \mp \frac{1}{2}\right) \delta(\mathbf{k} - \mathbf{k'} \pm \mathbf{q})$$

• The scattering rate, in the elastic and the equipartition approximation, is then of the form;

$$\Gamma_{k} = \frac{m^{*} k_{B} T}{4\pi \hbar^{3} k \rho} \left(\frac{e e_{\rho z}}{\varepsilon_{\infty} v_{s}}\right)^{2} \ln \left(1 + 4 \frac{k^{2}}{q_{D}^{2}}\right)^{2}$$

where  $q_{\rm D}$  is the screening wavevector.











### 3.2 Generation of Random Flight Times

• The probability of an electron scattering in a small time interval dt is  $\Gamma(\mathbf{k}) dt$ , where  $\Gamma(\mathbf{k})$  is the total transition rate per unit time. Time dependence originates from the change in  $\mathbf{k}(t)$  during acceleration by external forces

 $\mathbf{k}(t) = \mathbf{k}(0) - \mathbf{e}(\mathbf{E} + \mathbf{v} \times \mathbf{B})t/\hbar$ 

where **v** is the velocity of the particle.

• The probability that an electron has not scattered after scattering at t = 0 is:

$$P_n(t) = e^{-\int_0^t dt' \Gamma(\mathbf{k}(t'))}$$

• The probability therefore that an electron will suffer its next collision during dt around t is

$$P(t)dt = \Gamma(\mathbf{k}(t))e^{-\int_{0}^{t} dt' \Gamma(\mathbf{k}(t'))} dt$$

P(t) represents a non-uniform distribution of free flight times over a semiinfinite interval 0 to ... We want to sample random flight times from this nonuniform distribution using uniformly distributed random numbers over the interval 0 to 1, corresponding to typical numerical random number generators.

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 To calculate the carrier free-flight time, it is necessary to generate random numbers x with a given probability distribution f(x) over an interval (a,b) from evenly distributed numbers r.

(A) Direct Technique

• If  $P(\mathbf{r})$  is a uniform distribution between 0 and 1 then:

$$r = \int_{0}^{r} dr' P(r') = F = \int_{a}^{x_{i}} f(x) dx / \int_{a}^{b} f(x) dx$$

.

where  $x_r$  is a random number sampled from f(x).  $x_r$  is found by inverting this integration.

• Example, for constant f(x) is given below:

$$r = (x_r - a)/(b - a)$$
 or  $x_r = a + r(b - a)$ 





### (D) Self-Scattering

- The use of the full integral form of the free-flight probability density function is tedious (unless **k** is invariant during the free flight).
- The introduction of *self-scattering* (Rees, *J. Phys. Chem. Solids* **30**, 643, 1969) simplifies the procedure considerably.
- The properties of the self-scattering mechanism are that it does not change either the energy or the momentum of the particle.
- The self-scattering rate adjusts itself in time so that the total scattering rate is constant. Under these circumstances, one has that:

$$\Gamma = \Gamma(\mathbf{k}(t)) + \Gamma_{self}(\mathbf{k}(t)) \qquad P(t)dt = \Gamma e^{-\int dt \Gamma} dt = \Gamma e^{-\Gamma t} dt$$

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• Random flight times *t<sub>r</sub>* may be generated from *P*(*t*) above using the direct method to get:

$$r = e^{-\Gamma t_r}$$
  $t_r = -\frac{1}{\Gamma} \ln(1-r) = -\frac{1}{\Gamma} \ln(r)$ 

where *r* is a uniform random between 0 and 1 (and therefore *r* and 1-r are the same).

- Γ must be chosen (a priori) such that Γ> Γ(k(t)) during the entire flight.
- Choosing a new t<sub>r</sub> after every collision generates a random walk in k-space over which statistics concerning the occupancy of the various states k are collected.











# **3.5 Ensemble Monte Carlo Simulation** • For stationary problems, a single electron may be followed and statistics collected from time averages of the particle motion. • Estimators may be derived for the average drift velocity, energy, and particle distribution function (e.g.): $\overline{V}_{z} \cong \frac{1}{N_{s}} \sum_{i=1}^{N_{s}} V_{z}(t_{i})$ $S = \frac{\sigma}{\sqrt{N_{s}}}$ where $t_{i}$ is the *i*th free flight and *s* is the *standard error*, with $\sigma^{2}$ the variance of $v_{z}$ estimated by: $\sigma^{2} \cong \frac{N_{s}}{N_{s}-1} \left\{ \frac{1}{N_{s}} \sum_{i=1}^{N_{s}} V_{z}^{2}(t_{i}) - \left(\frac{1}{N_{s}} \sum_{i=1}^{N_{s}} V_{z}(t_{i})\right)^{2} \right\}$



 Estimates of the observable quantities are given by averages over the ensemble of particles at each time step (or multiples thereof)

$$\overline{v}_{z}(n\Delta t) \cong \frac{1}{N} \sum_{j=1}^{N} v_{z}^{j}(n\Delta t); \quad s = \frac{\sigma}{\sqrt{N}}$$

with the variance at each time step estimated as

$$\sigma^{2} \cong \frac{N}{N-1} \left\{ \frac{1}{N} \sum_{j=1}^{N} \left( \boldsymbol{v}_{z}^{j} \right)^{2} - \overline{\boldsymbol{v}}_{z}^{2} \right\}$$

• In steady-state, both ensemble and time averaging may be used to further reduce the error, and this is usually done in practice.













### 3.8 Carrier-Carrier Scattering

• For two-particle interactions, the electron-electron (hole-hole, electron-hole) scattering rate may be treated as a screened Coulomb interaction (impurity scattering in a relative coordinate system). The total scattering rate depends on the instantane-ous distribution function, and is of the form:

$$\Gamma_{ee}(\mathbf{k}_{0}) = \frac{m_{n}e^{4}}{\hbar^{3}V\epsilon_{\infty}^{2}}\sum_{\mathbf{k}}f(\mathbf{k})\frac{|\mathbf{k}-\mathbf{k}_{0}|}{\beta^{2}(|\mathbf{k}-\mathbf{k}_{0}|^{2}+\beta^{2})}$$
  
$$\beta = \text{Screening constant}$$

There are three methods commonly used for the treatment of the electron-electron interaction:

- A. Method due to Lugli and Ferry
- B. Rejection algorithm
- C. Real-space molecular dynamics

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### (A) Method due to Lugli and Ferry

- This method starts form the assumption that the sum over the distribution function is simply an ensemble average of a given quantity.
- In other words, the scattering rate is defined to be of the form:

$$\Gamma_{ee}(\mathbf{k}_0) = \frac{nm_n e^4 L_D^2}{4\pi \hbar^3 \varepsilon_{\infty}^2} \sum_{i=1}^N \frac{|\mathbf{k} - \mathbf{k}_i|}{|\mathbf{k} - \mathbf{k}_i|^2 + 1/L_D^2}$$

• The advantages of this method are:

1. The scattering rate does not require any assumption on the form of the distribution function

- 2. The method is not limited to steady-state situations, but it is also applicable for transient phenomena, such as femtosecond laser excitations
- The main limitation of the method is the computational cost, since it involves 3D sums over all carriers and the rate depends on **k** rather on its magnitude.

### (B) Rejection algorithm

• Within this algorithm, a self-scattering mechanism, internal to the interparticle scattering is introduced by the following substitution:

$$\frac{\left|\mathbf{k}-\mathbf{k}_{0}\right|}{\left|\mathbf{k}-\mathbf{k}_{0}\right|^{2}+1/L_{D}^{2}} \rightarrow \frac{1}{2L_{D}}$$

- When carrier-carrier collision is selected, a counterpart electron is chosen at random from the ensemble.
- Internal rejection is performed by comparing the random number with:

$$\frac{|\mathbf{k} - \mathbf{k}_0|}{|\mathbf{k} - \mathbf{k}_0|^2 + 1/L_D^2}$$

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• If the collision is accepted, then the final state is calculated using:

$$\cos \theta_r = 1 - \frac{2r}{1 + g^2(1 - r)L_D^2}$$
, where  $\theta_r = angle(\mathbf{g}, \mathbf{g}')$ 

where:

$$g = k - k_0; g' = k' - k_0'$$

The azimuthal angle is then taken at random between 0 and  $2\pi$ .

• The final states of the two particles are then calculated using:

$$\mathbf{k}_{0}^{'} = \mathbf{k}_{0} - \frac{1}{2}(\mathbf{g}^{'}-\mathbf{g})$$
$$\mathbf{k}^{'} = \mathbf{k}_{0} + \frac{1}{2}(\mathbf{g}^{'}-\mathbf{g})$$



 Using Newtonian kinematics, the real-space trajectories of each particle are represented as:

$$\mathbf{r}(t + \Delta t) = \mathbf{r}(t) + \mathbf{v}\Delta t + \frac{1}{2}\frac{\mathbf{F}(t)}{m^*}\Delta t^2$$

and:

$$\mathbf{v}(t+\Delta t) = \mathbf{v}(t) + \frac{\mathbf{F}(t)}{m^*} \Delta t$$

Here, F(t) is the force arising from the applied field as well as that of the Coulomb interaction:

$$\mathbf{F}(t) = q \left[ \mathbf{E} - \sum_{i} \nabla \varphi(\mathbf{r}(t)_{i}) \right]$$

• The contributions due to the periodic replication of the particles inside *V* in cells outside is represented with the Ewald sum:

$$\mathbf{F}(t) = -\frac{\boldsymbol{e}^2}{4\pi\varepsilon} \sum_{i=1}^{N} \left( \frac{1}{\mathbf{r}_i^2} \mathbf{a}_i + \frac{2\pi}{3V} \mathbf{r}_i \right)$$







- The particles are accelerated by the interpolated forces derived from the solution of Poisson's equation from the previous time-step.
- The particle-mesh coupling scheme consists of the following steps:
  - Assign charge to the Poisson solver mesh
  - Solve Poisson's equation for  $V(\mathbf{r})$
  - Calculate the force  $qE(r) = -q\nabla V(r)$ , and interpolate it to the particle locations (*q* is the charge of the particle)
  - Solve the equations of motion:

$$\frac{d\mathbf{r}}{dt} = \frac{1}{\hbar} \nabla_{\mathbf{k}} E(\mathbf{k}(t)); \quad \frac{d\mathbf{k}}{dt} = \frac{q\mathbf{E}(\mathbf{r})}{\hbar}$$





• In general, the charge on the  $p_{th}$  grid point is written as:  $\rho(\mathbf{r}_p) = \frac{q}{V_{cell}} \sum_{i=1}^{N} W(\mathbf{r}_i - \mathbf{r}_p)$ where  $\mathbf{r}_i$  is the position of the *i*th particle,  $V_{cell}$  is the volume of the  $p_{th}$  cell, q is the super-particle charge, and W is the weight function. • In the NGP scheme, the weighting function is of the following form:  $W(\mathbf{r}) = \begin{cases} 1 & -H/2 \le \mathbf{r} < H/2 \\ 0 & otherwise \end{cases}$ where:  $\mathbf{H} = (H_1, H_2, H_3) \longrightarrow \text{Vector describing cell dimensions}$ 



• When the Poisson equation is solved, the force acting on the *i*-th particle in the *p*-th cell is given by:  $\mathbf{F}(\mathbf{r}_i) = q \sum_p W(\mathbf{r}_i - \mathbf{r}_p) \mathbf{E}(\mathbf{r}_p)$ where the electric field at the *p*-th grid point is:

$$\mathbf{E}_{\mathbf{p}} = -\frac{1}{2} \left[ \frac{V_{p+1} - V_p}{\Delta x_p^e} + \frac{V_p - V_{p-1}}{\Delta x_p^w} \right]$$

$$\underbrace{\frac{\Delta x_p^w}{\Delta x_p^e} - \frac{\Delta x_p^e}{\Delta x_p^e}}_{X_{p-1} - X_p - X_{p+1}}$$
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## 4.2 Motion in Real Space

• The position in real space between collisions is determined by the coupled set of equations:

$$\frac{d\mathbf{r}}{dt} = \frac{1}{\hbar} \nabla_{\mathbf{k}} \boldsymbol{E}(\mathbf{k}(t)) \quad \frac{d\mathbf{k}}{dt} = \frac{q\mathbf{E}(\mathbf{r})}{\hbar}$$

• In the full band case, they are solved via Runge-Kutta numerical integration<sup>1</sup>. For parabolic bands:

$$\mathbf{v} = \frac{d\mathbf{r}}{dt} = \frac{\hbar\mathbf{k}}{m^*} \quad \frac{d\mathbf{k}}{dt} = \frac{q\mathbf{E}(\mathbf{r})}{\hbar}$$
$$x(t) = x(0) + v_x(0)t + \frac{q\mathbf{E}_x^0 t^2}{2m^*}$$

<sup>1</sup>S.E. Laux and M.Fischetti, in *Monte Carlo Device Simulation..*(ed. K. Hess, Kluwer 1991), 1-27.































